APPENDIX

TOTHE

DESCRIPTION and USE

OF THE

GLOBES.

CONTAINING,

- I. ANIMADVERSIONS upon the Structure, Apparatus, &c. of a Set of New Globes lately published; as also on the Book of their Uses.
- II. A New Construction of ORRERIES, with the great Variety of Phænomena they represent, and the extreme Degree of Exactness at a small Expence.
- III. Mr. GRAHAM'S New Inftrument for finding the LATI. TUDE at SEA by two Observations on the Altitude of the SUN OF STARS.
- IV. The Use of the GLOBE in finding the true DISTANCE of the Moon from the Sun | viewers Remarks.

- or a Star, by the observed Dis-
- V. The Rationale of correcting the PLACES of the STARS on the Celetial GLOBE when neceffary.
- VI. The Principles of the ancient Canicular ASTRONOMY explained by THEORY and Calcula-
- VII. An ABSTRACT of Mr. HORNSBY's Account of the enfuing TRANSIT of VENUS; with Directions subjoined for making a HELIOSCOPE to view the Phases to the best Advantage.

VIII. POSTSCRIPT to the Re-

The Whole illustrated by a large COPPER-PLATE.

By BENJ. MARTIN.

LONDON:

Printed for, and Sold by the AUTHOR, in Fleet-firest, and by the Bookfellers in City and Country.

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Ang Twenting the Arte work Roas Leanth and pulls occasioned by a late Publication of Mark Ciches, astronomy and a relative to tele defective in their Contract on left me wifted and the sale commented to sold the sale Photographic of the March and They are a more bary and from Wanner than it artifolog of residen Granted N. Comments of the Comment Find The state of the same of the same And the contractive of the contraction of the contr and takes the total of the The same of the sa and the state of t The party to the territory of the territory of the territory the said the state of the state of the end The state of the s Charles and Hall of the Control of the

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PREFACE.

Y writing this APPENDIX was principally occasioned by a late Publication of New Globes, which were pretended to be less desective in their Construction, less difficult in their Use, and contrived to solve the various Phænomena of the Earth and Heavens, in a more easy and natural Manner than any hitherto published.

If this had been really the Case, I knew what Sort of Fate must attend Mr. Senex's Globes, and consequently how far my own Interest would be thereby affected, as it is a principal Part of my Business to make those (hitherto esteemed) most compleat, most elegant, and useful Globes.

I judged it necessary therefore, to be well informed of the true State of the Cafe, and accordingly I purchased a Pair of the New Globes 12 Inches in Diameter. I then carefully examined their New Construction, and according to the best Judgment I am able to form of it, I think it so far from being any Improvement of the Globes, that it is quite the Reverse, viz. that it is a more defective Construction, more difficult to use, and much less adapted to explain the Phænomena of the Earth and Heavens than any Globes of the usual Construction. And Prudence, as well as Justice, require that the Particulars of this New Construction should also be laid before the Public, that they may be in some Measure able to judge for themselves; and fince, in all such Cases, no Person concerned should insift on his bare Word being taken, I am ready to satisfy any Gentleman of the Truth

of all I have advanced, by actual Inspection of the Globes themselves.

These are not my Sentiments of the Globes. only; I have found all that I have shew them to, of the same Opinion; and I can name two Persons of unquestionable Veracity, who will testify that a certain Gentleman of the ROYAL SOCIETY, and whose Name is well known in the Astronomical World, did most solemnly declare in their Hearing, that these New Globes were the WORST OF ANY hitherto made. After this, I need fay no more.

If any think I have treated the Author with less Ceremony than be deserves, they must give me Leave to inform them, that he must be conscious, none at all is due to bim from me. If it be faid the Stile is ludicrous and low; I answer, it is, and must be so, to be suited to the Subject; for whether you regard the Globes, or the Book-Spectatum admissi, Risum teneatis Amici?

Mr. Senex's Globes bave been fold in Preference to the New Globes, to those who have feen both, more than once; and they will never lose the high Estimation they have acquired

all over the World by the sensible and judicious

Part of Mankind.

I have added many other Sections upon Subjects very interesting and curious, and tend much towards extending the Use of the GLOBES, OR-RERIES, &c. as the Reader may be convinced of by perufing them.

Sept. 1, 1766. Fleet-Street,





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Animadversions upon the New Construction of New Globes lately published, and a Treatise of their Use, humbly submitted to the Consideration of the Public.

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lately published, has afferted, that their Construction is new and peculiar, and that they are contrived to solve the variaus Phenomena of the Earth and Heavens in a more easy and natural Manner, than any hitherto published—The first Part of this Affertion I readily grant is true, at least, as to their being peculiar; for that they have many Pecularities no Body, who has seen them, will pretend to dispute; some of B

which will appear considerable enough to

merit our peculiar Regard.

As to their being New, in Respect of their Structure, Engraving, and Manner of using them, it is allowed; and I grudge not the Author the extraordinary Glee with which he feems to enjoy this Novelty. They may also be said to be New in Regard to their Publication; but if we confider the Word New as opposed to Old, their Novelty in this Respect will admit of a Query. few Years can make no great Difference for a Terrestrial Globe, but the Case is quite otherwise in the Celestial One. By a New Celestial Globe, the public might naturally expect a Globe just now made, and the Stars all in their true Places, answering to their Places in the Heavens for the present Time. But here they must be aware of Mistakes; it being many Years fince these Globes were first put in Hand; and fince in 72 Years the Places of the Stars in Longitude and Right Ascension vary near one Degree, they will be out 30 Minutes in 36 Years; and 15 Minutes in 18 Years; and therefore will stand in need of Correction by Tables, when great Accuracy is required.

Hence Novelty can be no great Recommendation to a Celestial Globe in general, and never can give (of itself) a Preference to any particular Globe, by whomsoever made. For when Correction becomes necessary, as in a few Years it must, it may be as well made for one Globe as for another. But unless Celestial Globes are of a large Size, even this very Correction will be a vain Thing to propose. I am inclined to think that those who are competent Judges will see but little in this Argument——Few being weak enough to be taken with the bare. Bait of Novelty.

But these Globes (we are told) are peculiar, as well as New. Without any Criticism on the Phrase, I freely allow them to be possessed of the following Pecularities,

viz.

First, They are destitute of Hour Circles about the North Poles; and are the first I have ever seen or heard of, that are so. I believe very sew Globe-Makers are to be named, who did not know the Use of the Equator and Meridian, with Respect to the Division and Measure of Time; but not one of them all had the Ingenuity to demonstrate that they could with equal or greater Ease and Utility be substituted in the Room of the Horary Circle and its Index; the Students of Globes have therefore hitherto been deprived of the new and peculiar Privileges which THEY enjoy who purchase and practice by these peculiar Globes.

Secondly, The next Thing peculiar to these Globes, is, that they are fastened within their Frames, and in such a peculiar Manner

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also,

also, as renders it to no Purpose for any Person who is not a Mechanic to attempt to take
them out: The Wisdom of all preceding
Globe-makers has judged it absolutely necessary
that Globes should be at Liberty to be taken
out of their Frames, to answer many Purposes
which Use and Curiosity frequently suggest
to the Student—But by this New Invention it particularly appears, that every Gene-

ration grows wifer and wifer.

Third; The third Pecularity of the New Globes is in Consequence of the other two, viz. a Brass Semi-circle of round Wire, passing through the strong Brass Meridian, in which it is fixed by a Screw, as at each End it is fixed to the Wooden Circle of the Horizon; upon this, on each Side the Meridian, slides an Index or Minute Pointer. This Semi-circle not only shackles the Globe, but is productive of another Effect, such as no Man who makes, or uses Globes, would have thought of, or wished for, but the Author of these New Globes, who will securely enjoy the Honour of this Invention without Envy, viz.

Fourthly; It is peculiar to these New Globes that neither Pole can be elevated 90 Degrees, or raised from the Horizon to the Zenith; this we shall shew by-and-by, when we take Notice of the Accuracy of

these New Globes,

Fifthly;

Fifthly; Though the moveable Meridian is not peculiar to the Globes in Question, yet the placing a moveable Sun upon it on the Celestial, and a moveable or sensible Horizon on the Terrestrial Globe is really a Peculiarity; and by those whose Understandings are destined to be improved by these New Globes, will be found a very sensible Invention, and every Way worthy of the Author.

Sixthly; The Peculiarity which we are next pleasured withal, is an artificial Moon sliding upon a Silken String. If this be one of the humble Labours by which the Author has endeavoured to improve the Instruments of Science, for which he has boldly solicited the Patronage of his Majesty, I will venture to say, that were his Majesty's Patience not equal to his Goodness, he could never look upon this as an Object worth his Royal

Regard in the Use of the Globe.

Seventhly; There are many Arabic Words in Arabic Characters, engraven upon the Celeftial Globe, for the Sake of his Arabic Students, undoubtedly; that no one else can be benefitted by it, is evident, as there is no Interpretation on the Globe, nor any in the Book: But there is great Honour in being reputed a Linguist; and who knows but that the Author of these New Globes may be as well skilled in the Oriental Languages, as he is in the Mathematical Sciences. However this may be, it cannot be denied that Arabic Cha-

Characters on English Globes are both a Novelty and Peculiarity. And 'tis pity the Book wants that Decoration, for Want of an

Arabian Type.

Eighthly; The Manazil al Kamer, or 28 Mansions of the Moon, being placed all round the Zodiac, about and upon the Ecliptic, and Equator, must not only be esteemed an Improvement of this Instrument of Science, but the greatest Step towards an universal Eclaircissement in the polite Study of the Use of the Globes, and tend greatly to facilitate the Attainment of Astronomical Knowledge; and particularly in the New Method of discovering the Longitude at Sea. This is, indeed, a New Discovery of the Use of the Manazil; and their Insertion a singular Peculiarity of this New Globe.

Ninthly; The Addition of fourteen New Constellations, is (and I believe always will be) a Peculiarity in this New Celestial Glabe. The Sieur Desnos has done the same abroad; and I a little wonder at it in him, as I cannot persuade myself to think Mr. Senex (had he been living) would have done any such Thing. Indeed He added two trifling Ones, viz. Cor Caroli, and Robur Carol. but this was in Compliment to a King; and did no Harm to the Globe: But that He would have put on his Celestial Globe 14 Constellations, repleat with the smallest Stars, with the most vulgar and cumbersome Sym-

Symbols, and of Course not only useless in themselves, but prejudicing the Use of others, is not at all probable. Of the Ten Thousand Stars observed by De la CAILLE, he himself has made use of but very sew in his Astronomiæ Fundamenta, that were not known before: These Stars in the Firmament shew the Handy Work of their Maker, but on the Globe they shew just the contrary, in my

humble Opinion.

There are feveral other Things which the Maker of these New Globes pretends are new and peculiar, but some of them are too trifling to be mentioned; while others are afferted to be fuch with as little Truth it feems as he has Knowledge of the Matter. For who but a Person most miserably unskilled in the Art and Mystery of Globe-Making could ever affert so gross a FALSEHOOD, as that the Elevation of the South Pole was a Thing IM-PRACTICABLE in the Use of common Globes? 1 have not only made them, to have each Pole elevated with equal Ease (and in different Ways) ever fince I have been in the Trade, but they were made so long before; and I have no Fear of being accused of an audacious or mendacious Affeveration, when I aver that the worst Method I have seen used by others, is far more useful and rational than this of bis, though so new and peculiar.

The Author of these New Globes is pleased further to affert—That they are con-

trived to folve the various Phenomena of the Earth and Heavens, in a more easy and natural Manner than any hitherto published. But from what has been observed in the last Paragraph, our Author ought not to think it very indecent in me, if I put a Query upon the Truth of this Affertion likewise. As to the EASINESS of their Use, above others, wherein does that confift? Must not these New fo-contrived Globes be turned about on their Axis? Be elevated and depressed in their Frames? Have their borary Index fet to the Time on the Hour Wire? The Quadrant of Altitude fixed to the Meridian? And is not all this to be done by the Person who uses them? Then pray what Labour, what Manual Operation, is there less in these, than in other Globes? If none (for none I have been able to find,) then where is the greater Ease in using them.

Then as to the Truth of the second Part, viz. shewing the Phenomena more naturally than any others, as this is a Matter of still greater Consequence, we ought to have had much better Proof of it, than a bare Ipse dixit; Wherein does the artificial Globe differ from the natural one (the EARTH) but in the Appurtenances of an Apparatus? Does not that Globe, therefore appear most natural, that is least incumbered and shackled with these Appurtenances of Circles, Strings, &c. &c. &c.? The Question then

is, which is the most simple Construction, that of Mr. Senex's Globes, or that of the Globes lately published? They who have not an Opportunity of seeing both, may give a pretty near Guess, by considering that the Price of the latter is 5 Guineas, and of the former only 3, of the 12 Inch Size.

By the Globe formed into a Tellurian, the Phænomena of the annual and diurnal Motions are most naturally and simply represented, even by the Consent of our learned Author (for I suppose he means this Instrument by what he calls a Tellarium;) and if so, then how is this done but by the Instrument's having a greater Semblance to Nature in a less Apparatus of artificial Contrivances?

The Method he has invented to supply the Want of a Tellarium (whatever is meant by it) is as high a Thing as any in his Book (Page 53, &c.) and what is greatly becoming the KING'S Azimuth Compass-Maker, is to inform his MAJESTY, that the Variation of the NEEDLE at LONDON is between 19 and 20 Degrees; when his MAJESTY, no doubt, can inform him, in return, that the Variation is between 20 and 21 Degrees.

We have taken Notice of these Globes, in what relates to their being New and Peculiar, Easy and Natural; but the greatest Point of all remains as yet untouched, viz.

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The TRUTH of these New Globes. I don't find the Author has any where directly afferted, or descanted upon this essential Property of Globes; however as he has thought fit to acquaint his Majesty that be has made these Globes less desective in their Construction, it is the same Thing as to say, they give Solutions to Problems with greater Truth; for the less desective any Machine is, the greater Persection and Truth we may expect from the Operations depending upon it. It will therefore be worth while to enquire how far these Globes may be allowed to vie with, or exceed those of Mr. Senex in Respect of Truth and Accuracy.

I. And in the first Place let us cast our Eyes upon the strong Brass Meridian, and we shall find that it is almost HALF an INCH more in Diameter than the Globe itself, (in the 12 Inch size) whereas Mr. Senex's Globes have their Meridians little more than the Tenth of an Inch from the Surface; and consequently as the General Meridian is to represent those that are really fixed upon the Surface, those on Mr. Senex's Globes are about four times nearer to the Surface, and therefore so much nearer to the Truth,

than the other.

II. The same Desiciency from Truth in these New Globes, and the same Proximity to it in those of Mr. Senex, appears also in the Hori-

HORIZON, or broad Paper Circle in our Author's New Stile.

III. Both the North and South Parts of the Horizon, which should have been (as in other Globes they are) contiguous to the graduated Side of the Meridian, are cut away in these New Globes to the Distance of full of an Inch: So that with Regard to all Problems relative to the LATITUDE of Places. DECLINATION of Stars. ELEVA-TION of the Poles, BEARING of Places, AMPLITUDES, AZIMUTHS, &c. of Stars, I say, in all these grand Problems at what an enormous Distance are we from the Truth! And therefore when these Globes. are afferted to be of a less defective Construction than others, we are quite at a loss to conceive what Opinion our Author could entertain of the Judgment or Sense of those, who were to use them.

IV. But if any Advocate can be found for any Thing so bad as the above Construction, he will put us in Mind of an Advertisement at the End of the Preface, wherein you are directed to remedy all these Inconveniences by cutting of a Card secundum Artem, and by applying it, aprepos, to the Surface of the Globe, Meridian, Horizon, &c. &c. &c.

But will every Purchaser of these Globes conceive he has Dexterity enough for executing such a nice Affair? Or if he has, will he not think it hard, that after having purchased.

chased Globes at an excessive Price, he should still be left to put his Ingenuity to the Rack to construct Materials and Implements for using them, himself? But supposing still further, that he has as much Patience as Money, yet how is he to posite the Card upon the broad Paper Circle, so that its Edge may be applied to the strong Brass Meridian in such Manner as that the Degree, and Parts of a Degree, may be ascertained with such sufficient Accuracy that the North Pole may thereby be elevated EXACTLY.

V. But let us hear what the Author of these less defective and less difficult Globes is pleased (or forced, shall I say?) to declare of them himself, these are his Words-" It is frequently required to know what Point upon the strong Brass Meridian or broad Paper Circle, exactly answers to a given Point upon the Globe, and (fayshe) this CANNOT BE WELL KNOWN BY INSPECTION, ON ACCOUNT OF THE NECESSARY DIS-TANCE OF THESE TWO CIRCLES FROM THE SURFACE, &c. Here then is an ingenuous and bonest Confession of the plain Truth. Such CANDOUR is rare to be found in an Age abounding with fo much Cunning, and especially in TRADESMEN characterifing their Own New and peculiar PERFORMANCES.

VI. The next Proof of these Globes being less desective than others, is, to be sure, that the Equator and equinoctial Line is entirely eclipsed from our direct View by the Opake

Body

Body of a Brass Wire placed directly over it.

Sir Francis — who lately made a
Purchase of a Pair of these Globes, declares
he is obliged to bend his Head as obliquely
as a Parrot does, to get a View of this
important Circle, and its minute Divisions,
when he is to find the Hour and Minute,
Right Ascension, Oblique Ascension, Culmination, &c. of any of the heavenly Bodies,
which are all to be sought for in that Circle.
And Captain Surge is constantly execrating
this Brass Circle over the Equator, which
he believes will prevent (at least make it very
difficult,) for him to find the Longitude

by SEA or LAND either.

VII. Another Proof of these Globes being less defective than others, the same Brass Wire will very readily furnish; (of what prodigious Service is this Pittance of Wire upon these New Globes!) For being itself one Tenth of an Inch Thick, and carrying an Index, which is twice that Thickness, it must cause, that in the best Position of the Index, the equinoctial Line (or the Equator) cannot come within To of an Inch of the Horizon in the South or North Parts, which is a whole Degree; and if the Hour Index be placed near Aries or Libra, the faid Lines will want feveral Degrees of Coincidence with the Horizon, and by just so much, or so many Degrees will it be impossible to bring the Poles into the Zenith, or elevate them to 90 Degrees. VIII. Hence VIII. Hence it follows that only two of the three different Positions of the Sphere can be truly represented on these Globes; that of the parallel Sphere being impossible, where the Pole cannot be placed in the Zenith. And may we not deem this a most flagrant Proof of the less defective Construction of these New Globes?

IX. It will be found by measuring the QUADRANT of ALTITUDE, that it is a Quarter of a great Circle upon the Globe, and is accordingly divided as fuch, but when one End of this Quadrant is elevated and affixed to the Meridian at so great a Height above the Globe, and the other End strongly springing against the Horizon at such Distance from the Globe, will this New Globe-Maker fay, that this is in fuch Circumstances a Quadrant of Altitude, i. e. The proper . Measure of the ALTITUDES of the Sun or Stars above the Horizon, or of their AzI-MUTHS from the Meridians? Every TyRoin Spherics knows it becomes only a Segment of Circle of a large Sphere, improperly divided into oo Degrees; and can only touch the Globe, for which it is made in the Middle Part, instead of lying closely upon its Surface throughout. Lo, here, is another Instance of a less defective Construction of Globes!

X. But what will give the most exalted Idea of a Genius for Improvements is, that in Problem XV, (Page 68.) To find the Sun's

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Sun's Altitude, he has directed the artificial Sun to be placed over the Sun's Place in the Ecliptic, and then by moving the graduated Edge of the Quadrant to the Center of the artificial Sun, he tells us, the Degree of the Quadrant which is cut by the (artificial) Sun's Center, is the Sun's Height at that Time. But in that Set of 12 Inch Globes of this New Construction which I bought, when I was following these fage Directions, and had at last fumbled out a Solution of the Problem on this New Celeftial Globe, and also on that of Mr. Senex's for the same Moment of Time, I found the Height of the natural Sun upon the latter to be more than two Degrees less than that of the artificial Sun upon the former. Now 'tis evident that one of these Globes must be shamefully defective in Construction indeed; but which, is left to the Reader, partial, or impartial, to judge.

XI. From hence may be gathered the Reason, in general, why so very sew of the Problems in this extraordinary Book are illustrated by Examples; for had this 15th Problem been thus illumined, we had seen stoo clearly perhaps) the Defect of Construction in Comparison of others; and this would have been a woeful Affair after he had given his Majesty an Assurance of the Contrary. In sine, those who have to do with artificial Suns and Moons, ought to be content with artificial Altitudes, Azimuths, &c. But such

as aspire to nothing more than what is purely simple and natural, may find it all in Mr. Senex's Globes.

XII. The Method of engraving the Calendar and Degrees of the Sun's Declination (usually called the ANALEMMA) upon the Meridian about the North Pole is nothing New or peculiar, it being the Invention of the late Mr. Joseph Harris, about 30 Years ago, which was probably before these Globes were began. But Mr. Harris observes at the fame Time (N°. 456 of the Trans.) that GLOBE-MAKERS might fave the TROUBLE and Expence of engraving this upon Brass, by doing the fame Thing upon the Paper of the Globes; this Advice Mr. SENEX took; but they who pay 5 Guineas for Globes instead of 3, ought to have some valuable Confideration for it, viz. a BRASS ANALEMMA. instead of a Paltroon paper one, that will cost little or nothing.

XIII. The Crepuscular Circle is another Article, neither New nor peculiar; and I think moreover, if it does not prove the Construction of the Globe defective, it proves itself very defective; being only Scantlings of common Brass Wire of Three Half Penceper Oz. Those made by Mr. Senex were doubtless of a different Construction; and the Globes made since his Plates came into my Hands, have never been disgraced with rough and un-

wrought Materials, and so inartificially placed as to be of little or no Use.

XIV. In regard to the Compass, as it is, and always has been an Appendage to Globes, I have only this to observe, that Mr. Senex's Method of applying this useful Part, is greatly preferable to that we find in these New Globes; is much less expensive, and fixed

with greater Certainty and Truth.

XV. All the Articles hitherto mentioned are relative to Deficiency; let us next see what may be found redundant in these New Globes; and the first Thing observable of this Kind is not only New but very peculiar indeed; we are told by this Author (Page 43.) That there are Four Colures. I must confess this is the first Time I ever heard of that Number; Mr. Senex never had more than two upon his Globes: Nor can I find above THREE out of the Four upon these New Globes, viz. one Solftitial Colure, and two Equinoctial Colures; it is true, indeed, he speaks of Solstitial Colures in the plural (Page 44) and fays they are made out of two Meridians passing through the Points marked 5 and w; upon other Globes there is but one Meridian which can possibly pass through these two Points. If we allow of this new Invention for making Colures, we shall find at least five upon these New Globes, viz. two Solfitial, and three Equinoctial Colures. For befides the equinoctial Colure paffing through through v, and the other Equinoctial Colure paffing through a, there is it seems, a third Equinoctial Colure passing through the 24 Degrees of the Equator, East of the Point It is true, he does not (totidem Verbis) fay it is the EquinoEtial Colure, but only the Colure, or Equinoctial Meridian, A. D. 1753, o a. And this fifth Colure is diftant from v, or the first Meridian passing through Greenwich Observatory, 156 Degrees to the West; for the wonderful Invention of which, and the extraordinary Ufe made of it, fee Page 126, &c. of the Book, by which the Reader will be convinced that the faid Book is every Whit as new and peculiar as the Globes it describes; it being the first Time undoubtedly that ever an equinoctial Meridian was heard of that did not pals thro' the Equinoctial Points.

XVI. In other Globes it has been always adjudged sufficient to number the Degrees of Longitude each Way to 180 Degrees from the first Meridian; and all the modern Books, Maps, Tables, &c. are formed according to that Method of reckoning; but the Purchafers of Our New Globes are not to be sobb'd off with Things done thus to balves. He not only gives them the two Halves, but the Whole Circle into the Bargain, and roundly reckons on from 0, to 360 Degrees without stopping. The Equator thus surrounded by two Phalanges of Figures above, and one of a double

a double Notation below, appears numerically equipped beyond any Equator upon a

Globe, fince Eudoxus's Time.

XVII. I have already taken Notice of the great Addition made to the Number of the fixed Stars upon this New Celeftial Globe, which we are told are feveral Thousands more than have appeared upon other Globes. And as it must be allowed there is at least a Thoufand more upon other Globes than can be of Use to Mankind in any Shape or Sense whateever, to what Purpose is all this To-do about Stars? Are they intended to illumine the Globe? If fo, why have they fuch a caliginous Aspect on this New Globe! In the Heavens they appear bright and glorious; but I appeal to any One who views the Southern Hemisphere of this Globe, if it does not look more like Blotting Paper, than the Firmament of glowing Stars.?

XVIII. Under the last Head of Redundancies I may justly reckon the amazing Multiplicity of Names of Places upon the Terrestrial Globe, and the most Part of them of little or no Importance; but if the Peruser of such a Globe has no Objection to its being thus crouded, in a small Character, I have none, he may suppose. And if the Author had authenticated his Innovations upon the Terrestrial Globe with any more than his Ipse dixit, it might have prejudiced the Sale of Mr, Senen's; but as the Case stands,

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I believe they are pretty fafe, and so will remain 'till other Lights are thrown upon Geography than what can be expected from this

Quarter.

I have done with the Globes; and shall conclude with a few Remarks on the Book that describes them. As to the Grammatical Part of it, I will say it is much exceeding what I expected from an Adverisement of it three times repeated in the public Papers. I wonder none of his worthy Friends could be found to give him a Jog before. In Orthography he is not quite so correct as one might expect; for fure the hackney'd Word Phanomena is become too common to admit of the Excuse, Græcum est, non potest legi. However our Author is resolved to be right by Hook or by Crook, and therefore spells it every Way it can be spelt, viz. Phoenomena, Phanomena, and Phenomena. As to the Word Tellarium, I protest I can find it in no Dictionary, nor can I derive it from any One, or all the Languages together, I ever heard of. His Eskmios for Eskmaux, Kanton for Canton, Antego for Antigua, and a hundred others, will aftonish no Body that has read this peculiar Book.

In the Title-Page we are told One great Design of the Book is to shew the Cor-RESPONDENCY of the TWO SPHERES. And though this new Doctrine employs near half the Book, yet after looking it over, I was

just as wife as before; for after all the learned and tedious Labour he bestows upon us about the natural Agreement between Celestial and Terrestrial Spheres, Solar Correspondents, Celestial Correspondents, Correspondency of the fixed Stars, Signal or Warning Stars, and many Tables of Correspondency of Stars and Places, employing near 40 Pages of the Book. I was fo stupid as to see no more Correspondency between one and the other Sphere. than appeared before the Publication of these New Globes. Nay his Descant upon Signal or Warning Stars, and their Perpendicularity over Peoples' Heads, must be incomprehensible to those who have studied it no more than myself.

As to the Phænomena of the HARVEST-MOON, we find nothing New in the Book; and the New Globes will explain them in no peculiar Manner, not even so easy and natural as may be shewn upon the common Sort of Globes, if we are obliged to use his artificial Sun, Moon, Wires, Strings, &c. which are surnished for such Kind of Pro-

blems.

For a Specimen of Astronomical Skill, that is certainly both new and peculiar, I must refer the Reader to the Book, where from Page 126 to 148, he will find the Method of computing the Time of the Autumnal Equinox, and the Difference of Time between that and the Vernal Equinox, with Precepts, Tables,

Tables, &c. of Retrocession, Autumnal Equinoxes, &c. with Calculations for Alexandria, Greenwich, &c. with a Circle for the literal Character of Days, &c. But because these Things do not concern the Globes, therefore the Reader must be left to find out the Meaning himself, and make the best Use of it he can.

The Methods of finding a Meridian Line, the Equation of Time, and shewing how the Terrestrial Globe will represent the real Phenomena relating to the Earth when actually compared with the refulgent Rays emitted from the great Sphere of Day, are as incomprehenfibly fine, as the Language in which they are presented to us; but here we must follow our Author's Example, and suppose, the many Advantages arifing from these Capital Problems relating to the placing the Globe in the Sun's Rays, an intelligent Reader will eafily difcern, and readily extend to his own as well as the Benefit of his PUPIL. N. B. By an intelligent Reader you are to understand a Teacher of the Use of the Globes.

I can't conclude without taking Notice of one very extraordinary Paragraph in the Preface, though it is purely Astronomical, and in which the Globes are no ways con-

cerned; it runs thus;

"The Table of the Passage of the first Point of Aries over the Meridian, is taken from an EPHEMERIS of (De) LA CAILLE, with with fome LITTLE ALTERATIONS; and has been carefully compared with twenty Years CALCULATIONS made by the fame Author. The Necessity, fays he, I found myfelf under of haftening the PUBLICATION of THIS TREATISE, and the various Interruption and Avocations from which a Person in MY STATION cannot be exempt, would not allow ME sufficient Leisure to CACULATE an ORIGINAL TABLE." Here the Public has a right to propose the following Queries. viz. (1.) Why any one has prefumed to make any Alterations in Tables of any Sort calculated by the late DE LA CAILLE, without giving a Reason for his so doing? (2.) How have these Tables been carefully compared with 20 Years CALCULATIONS? (3.) How does it appear that the most able Astronomer which this Age has produced, both for Calculations and Observations, has offered Us a defective Table of the Passage of the first Point of Aries over the Meridian? (4.) and in what Ephemeris of DE LA CAILLE is there any fuch Table to be found? (5.) In his Ephemerides of the Celestial Motions for ten Years, viz. from 1765 to 1775 there is a Table of the Distance of the first Point of Aries from the Meridian of Paris, and is variable thro' every Year. Can this be the Table in Question? (6.) Whatever Table is here referred to, and whatever be its Deficiency, if it was of DE LA CAILLE's compiling, have we any Reason to expect it to be better done de novo by our Author from any Specimens we have of his astronomical Abilities contained in this wonderful Work?——
From the little I know, or can judge of astronomical Matters, I could not have supposed there was a Man living who would have presumed himself able to mend any such Table of the Abbe De la CAILLE.

Our Author has great Hopes that the Superior Accuracy with which the Plates are drawn and engraved will appear at first Sight to competent Judges; but if the Public were certified of the Truth of what I have heard from the Workmen, viz. that the Plates were done over and over again, in Whole or in Part, and many Changes and Sets of Hands fucceffively employed for many Years past; and that at last it would have been impossible to have got them compleated, had they not hit upon the Expedient of copying Mr. Senex's Papers of the 17 Inch Celestial Globe. I fay if they believe these common Reports, there will be very few competent Judges found to have any extraordinary Opinion of the superior Accuracy of copyed Plates, or to discern it at first or second Sight either.

As to the Drawing and Engraving of the Plates of the Celestial and Terrestrial Globe in the Book, they are equally incomparable and peculiar, and quite beyond the Reach of

Criticism. And since they are said to represent these Globes AS they are improved and constructed by the Author, it will save us the Trouble of saying any more about the Globes,

Plates, or their Engraving.

K. Solft J.

I shall dismiss this Subject with only one Reflection on our Author's Talent at moralizing upon the Works of Nature; for speaking of the PLANETS he observes very gravely, and with great Peculiarity, -That the Regularity of their Motions, STRICTLY CON-FORMABLE at all Times to the LAWS of THEIR CREATOR, exhibits a STIKING PAT-TERN of OBEDIENCE to every RATIONAL Spectator. But why are we fent to the Planets to learn this Piece of Rational Deportment, fince every Man's CLOCK affords him the same striking Patern of Obedience. But if our Author thinks Rational Beings will admire his mechanical Methods of becoming devout and religious, it is no Wonder if he imagines that his Globes will open a large Field of Geographical and Astronomical Knowledge, fraught with Instruction and Amusement; fince these Things naturally flow from mechanical Constructions and Contrivances; whereas the Principles of Virtue and Religion, in order to their being rational, were never before known to be founded in Laws of Nature designed only for inanimate Beings, and, of Course, are purely Mechanical. CASHINA TO JAN STREET E

In thort the Reader of this Book ought to have a new and peculiar Understanding; how else will he comprehend how a Circle may represent a Plane (Page 9)? Or how to supply the Want of a Tellarium (Page 53)? Or how the Equator can (in these Globes) ever be in the Plane of the broad Paper Circle (Page 60)? Or what is the nearest Mean Length of a Tropical Year (Page 127)? Or what is the Use of the Manazil al Kamer (Page 205)? Or what is the Shade of Extuberancy (Page 224)? Or what are the Four Cardinal Points of the DAY? Or what the Use of the Tables at the End of the Book? To pass by many other Things equally incomprehenfible by vulgar Capacities.

The Undertaker of a Talk fo complicated and laborious (as writing a Treatise on the Use of the Globes) hopes the Reader will make some favourable allowance for him. and correct the Errors for himself, if there should appear to be any. This is but a reafonable Request, and it would be cruel to deny it; Sins of Ignorance (when we are not needleffly guilty of them) are always venial; fince bumanum est errare. But why, fince the Task is in itself so grieviously hard, should the Undertaker, take so much more Pains than was necessary? Who besides this Undertaker would undertake to use the Periphrasis, OUR NEW GLOBES 25 times over, when so many fingle Words in other Writers fuffice ?

[27]

fuffice? Or the Circumlocution STRONG BRASS MERIDIAN, no less than 100 Times? And theirksome descriptive Tautology, BROAD PAPER CIRCLE (for the Horizon) 75 times repeated? Indeed when we take a proper View of this Work we cannot but exclaim with the Poet, Hic Labor, boc

Opus!

I don't apprehend there is any more Complement than Truth, in the Affertion, That it is the Privilege of real Greatness not to be afraid of Diminution by condescending to the NOTICE of LITTLE THINGS -- Whatever our Author may presume on that Supposition, I believe most of his MAJESTY's literary Subjects have a better Opinion of his MA-JESTY'S Judgment than to suppose he can look upon any Thing here called bumble Labours in any other Light than that in which the Author himself has truly placed them, viz. Of LITTLE THINGS, indeed! And whether his MAJESTY's literary Greatness may be liable to no Diminution by condescending to the Notice of them, must be left to the Determination of his most indicious and impartial Subjects.

Thus we have considered most of the Phoenomena of a Book dedicated at first to a King, and may at last perhaps, have the Honour of being devoted to a Goddess, notwithstanding all the Correspondency of all

the Spheres it talks so highly of.

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At

At the End of the Book we find a CATA-LOGUE OF INSTRUMENTS, prefaced in a STRAIN of RHETORIC that is also New and Peculiar. But what I shall take Notice of are the following Expressions, viz. such Instruments as are either invented or improved by bimself, and constructed according to the MOST PERFECT THEORY. Now I believe I may fafely appeal to the Public, and ask if they do not remember those very Words in my Advertisements in the public Papers for many Years past? However, any Gentleman may fee them in my Shop-Bills at any Time. Now he either stole them from me, or I stole them from him; but this Matter will not long be a Subject of Dispute, for unless he can produce Bills or Catalogues which have those Expressions prior to mine, I shall charge him with the Theft: He tells us, he would not willingly incur the Imputation either of PLAGI-ARISM or INGRATITUDE; but unless he can acquit himself of the above Charge, he must appear guilty of Plagiarism; and also of Ingratitude of Course, because he took it without thanking ME for it. But how much baser must that Disposition still be, which can permit a Man to take any Thing from another, to which he knows he has not only no Right. but is conscious his Behaviour for Ten Years past has rendered him an Object of just Contempt to the Person he steals from? There are other Articles of Impeachment, such as Achromatic Refractor, Manual Orreries, Planetariums, and I suppose his Tellarium, are all taken from my Bills, &c.

But when he reprints his Catalogue I must beg the Favour of him not to disgrace my Instruments with such false Spelling, but write it Tellurian, as I have done. Also some common Words would better become his Catalogue, if they were more properly and intelligibly spelt, as Pantographer, Zograscope, &c. This last Word Zograscope is entirely new to me, and I believe peculiar to this Author alone; the same may be said of the Word Logarithim in some other Writings of his Majesty's Mathematical Instrument-Maker.

As we are told, that his Instruments are constructed according to the MOST PERFECT THEORY, it would have been but becoming in him to have first convinced the Public that he knows when they are so, and when they are not; for we are not to take a Man's Word only for a Proof of his being a Mathematician; besides the Assertion is not true in his Mouth, for he never constructed some Instruments according to the most perfect Theory; for Instance, this requires a different Construction and Application of the Line of Latitudes, than what you find upon any Scales of his Make, unless he has stolen that likewise from my

Dialling-Sector. And were he to fell no more instruments than what he could demonstrate the THEORY of from genuine PRINCIPLES of GEOMETRY, I believe it would not be the worse for the Rest of the Trade.

He has a New and Curious Hydrometer in Hand, it seems; and it very luckily happens, that he could not have pitched upon such another Instrument for giving him an Opportunity to exert his utmost Faculties, and whose Theory displays so ample a Field. The Truth of which I have found by seven Years Experience; the Physical and Mathematical Principles of the Construction of my Hydrometer, I have long since published; And its hoped my Antagonist will not fail to do the same by his; The Merits of each will then be easily decided by competent Judges among the Honourable Commissioners of his Manuerry's Customs, &c.

I should have taken in very good Part all his Plagiarism and subreptitious Methods of Dealing by me, if he had not been so ungenerous as to call me Names for it when he had done. He would have been welcome to all the Eclat my New AIR-Pumps have added to his Shop-Windows for many Years past, had he been genteel enough to have told his Customers who they were indebted to for them. No Instrument invented by him, ever appeared in my Windows; Not but that

that if he should at any Time oblige the Public with any Thing that is really better than what they have already, I sincerely declare I shall be the first to acknowledge its Merit; and would scorn to depreciate any Work of public Utility, from what Hand soever it comes; for with me, Fas est et ab Hoste doceri.

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SECTION II.

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A New Construction of Orreries, flewing the Great Variety of Phænomena they represent, and the extreme Degree of Exactness they admit of at a small Expence.

N HAM in his Artificial Clock-Maker, has treated upon the Subject of Orreries, the Mechanism thereof, and the Truth of such Work. And what he has delivered is quite upon the Old Plan, and chiefly historical.

Nor has any foreign Author besides the Great Hugenius, touched upon Calculations of this Kind, that I know of. This noble Author, in his Tract called the Automaton, has delivered the Principles of Calculating Numbers for the Planetary Motions, and by which he has constructed his Automaton or Planetarium to such a Degree of Truth in the Mean Motions, Excentricities, &cc. as far exceeds any Thing of the like Sort done by any other Hand.

But,

But no human Productions are perfect; Hugenius leaves us to investigate Numbers for the Movement of the Moon's Node, and Apogee, the Mean Motions of the Satellites of Jupiter and Saturn, and the Phases of Saturn's Ring. But, what is the greatest Desiciency of all, there is not a Word of the Mechanism of the Earth's Diurnal Motion about its Axis.

It is this last Article, indeed, that distinguishes an Orrery from a Planetarium, which has nothing more than the Annual Motions. As Hugenius made use of Ricciolus's Tables, his Numbers and Constructions are not to be expected so correct as those deduced from modern Tables, viz. those of Halley, De la Caille, Mayer, and Clairaut.

From these most correct Tables I have been at no small Pains of deducing a System of Numbers, which will shew the Annual Mean Motions of the Planets; the Parallelism of the Earth's Axis; the Diurnal Motion of the Earth; the Menstrual Motion of the Earth and Moon; the Motion of the Moon's Node, and Apogee; the Motion of Jupiter's Satellites; the Motion of Saturn's Satellites; with the Parallelism of his Axis for the Phases of his Ring. All of which are therefore in as great Perfection as the present improved State of Astronomy will admit of.

But that the Reader may be in some Degree apprized of the exceeding Accuracy of these Numbers, or how near to the Truth they approach, I shall lay before him the

following Sketch for his Satisfaction.

The Planet MERCURY makes 847 Revolutions in 204 Years in this ORRERY; but in the Heavens, it revolves 847 times in 204,008 Years; this gives an Error of 16' 56" 16" in 20 Years; which is not one Minute per Annum.

The Planet VENUS in the ORRERY makes 395 Revolutions in 243 Years; but in the Heavens she revolves the same Number of times in 394,9965 Years; which gives an Error of 6' 28" 40" in 20 Years, or about 20"

per Annum.

The EARTH's Period about the Sun is

here stated at 365 5 49' 25".

The Planet MARS revolves in the ORRE-RY 67 Times in 126 Years; but in the Heavens the same Number of Revolutions is compleated in 66,9977 Years; hence the Error in 20 Years is not more than 7' 53" 24";

which is about 24" per Annum.

The Planet JUPITER makes 14 Revolutions in the ORRERY in 166 Years, in the Heavens it performs 14,0018, Revolutions in the same Time; the Difference gives 4' 40" 48" for the Error in 20 Years, not quite 12" per Annum.
F 2

The Planet SATURN revolves about the Sun in the ORRERY 7 times in 206 Years; but in the Heavens it makes but 6,99895 Revolutions in the fame Time. Whence Saturn is too fast by 1' 37" 12" in 20 Years, or 4" ½ per Annum.

For the LUNARIUM.

As the Time of Mean Synodical Revolutions of the Moon is 29^d 12^h 44' 2" 53", there will be 21038 fuch Lunations in 1701 Years, compleatly; in the Orrery there are 1781 Lunations in 144 Years, which is so very near the Truth, as to differ from it only 28' 30" in 20 Years, or but a little more than one Minute per Annum.

The Mean Annual Motion of the Node of the Moon's Orbit being 19° 19' 43", there will be 67 Revolutions of the Node in 1247 Years; and thence the Mechanism of the Motion of the Node in the Orrery will be

extremely near the Truth.

The Annual Mean Motion of the Moon's AFOGEE being 40° 39' 50", there will be 13 Revolutions in 115 Years; whence also this Motion is represented with great Accuracy in the Orrery.

There may also be added to this Orrery, a second LUNARIUM, for shewing the true Phases of the Moon and her Librations both in Longitude and Latitude; which Part has

not yet appeared in any Orreries that I have

The Parallelism of the Plane of the Ecliptic in this Lunarium is adjusted to the precise Truth, and the Days of the Moon's Age, and her Latitude in every Part of her Orbit for any given Time, are distinctly shewn.

The TELLURIAN.

In this Part our Numbers are so very exact, that the Axis of the Earth is not only always perfectly parallel to itself, but the Diurnal Motion of the Earth about its Axis is here adjusted to the Mean Motion of the Sun to so nice a Degree, that they both begin and end the Year together within less than 30 Minutes per Annum. See further concerning the Uses of this curious Part in the foregoing Treatise of the Globes, Page 168, &c. and in the Young Gentleman's and Lady's Philosophy in the General Magazine for the Year 1760.

The Jovian System.

In order to exhibit the Mean Motions of the four Safellites of Jufiter, we have only to compare their several Revolutions with that of the *Primum Mobile*, and from thence we obtain the Numbers for the Construction of this *Jovian System*, with all the Accuracy the present State of Astronomy will afford. These periodical Times and Numbers are contained in the following. Table.

Satell.	Periodical Times.	Revol.	Days.
T	D. H.	DID. I	
II.	1 18 27 35		34
iii.		20	71
IV.		116	267

The Saturnian System.

In the same Manner you observe the Periodical Times, and Ratios of the Numbers which compose the Mechanism of the Saturnian System, in the Table below.

Satell.	Periodical Times.	Revol.	Days.
1.	D H I 21 18 27	98	185
II.	2 17 41 22	19	. 52
III.	4 12 25 12	29	131
IV.	15 22 41 12	55	877
V .	79 7 47 0	34	2697

The Phases of Saturn's Ring are provided for by the proper Inclination and Parallelism of its Axis. So that at any Time of his Period they are the same in the Orrery, as in the Heavens; and thus all the Phanomena of the Annual, Menstrual, and Diurnal

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Diurnal Motions of the Planets, Primary and Secondary, are exhibited in this Construction of an Orrery with much less Expence, and far greater Accuracy, than could be attainable with Numbers less perfect than those we now have; not only so, but I have never seen, or heard of any Orrery, that will exhibit all the above-mentioned Appearances with any Numbers, or by any Construction at all, not even those which have been sold for 1000 Guineas.



SECTION. III.

Mr. GRAHAM'S New INSTRUMENT and METHOD of finding the Latitude at Sea, by two Observations of the Altitude of the Sun or Stars.

**HIS Problem so very difficult in Theory and Calculation, and yet so extremely necessary to the Perfection of Navigation, will admit of a very easy and exact Solution by the Globe, or an Instrument made of Part of a Globe, which was some Years ago described in

in the Philosophical Transactions, Vol. VIII. Page 371, by the Inventor Mr. R. GRAHAM; this Account of so useful an Instrument and Method of finding the Latitude we shall here transcribe, as it immediately depends upon the Nature and Use of the Globe, which, if it be of the largest Size, or that of 28 Inches in Diameter, will give a Solution to the Exactness of a Minute with great Ease. His Account is as follows, in his own Words.

The Necessity of finding the Latitude a Ship is in, is too well known to be infifted on: Frequent Opportunities of observing the Latitude must consequently be of very great Advantage to Navigation. The Method usually practifed, is by taking the Sun or Star's Meridian Attitude or Zenith Distance: In this Case, if the Sun does not shine but for fome fmall Time only, before Noon and after, though it be clear all the rest of the Day, it is of no use for this Purpose. Mr. Fatio. F. R. S. (in the Year 1728) proposed a Method for finding the Latitude, from two or more Observations of the Sun (or Stars) at any Time, the Distance of the said Observations in Time, being given by a Watch; but as his Method requires a vast Number of Computations, and a great deal of Skill in Spherical Trigonometry, it has very feldom been made wie of, and never but by good Mathematicians. The Influment here defcriber

ai

scribed will answer the same End, and has these Advantages, viz.

1st. It may be very eafily understood by

Sea-men.

2dly. It immediately shews the Latitude of the Place.

3dly. It gives the Time of Day at Sea, when no other Instrument can.

4thly. It may be made as large and conse-

quently as accurate as is defired.

A BC(Fig. 1.) reprefents part of the Hemisphere of a large Globe (half the Globe, and the Part below the Tropick are cut off, that it may take up the less Room). AC, half the Equator, divided into 12 Hours above, and 180 Degrees below, and subdivided into Minutes, as is likewise the lower Tropick DD; EE, a moveable graduated Meridian, turning on the Axis FF. G an Index to fix it (by the means of the Screw H) to any Hour. IiI, a circular beam Compass, the Center I i to be fixed on the Meridian to any Degree and Minute of Declination, by the Method commonly called Nonius's Divisions: k the Point for drawing Arches, which is likewise fixed to any Degree and Minute by the same Method. As the Meridian is at some Distance from the Globe, L is a piece of Brass to fix on the Meridian, marked with Nonius's Divisions, with a Point reaching down to the Interfection of the Arches, by which Means the Distance of the said Interfection

fection from the Equator, or it's Latitude is found. The Degrees and Minutes may likewife be shewn by diagonal Lines.

Prop. I. From two Observations of the Height of the Sun, the Distance of the said Observations in Time, being given by a Watch, as likewise the Declination of the Sun; to find the Latitude of the Place, and Hour of the Day.

I. When the Ship is at Rest, that is, at Anchor, or in a Calm, so as to have little or

no Progressive Motion.

Case I. Suppose the Sun in the Equator, on the Day of Observation: Fix the Centre of the Beam-Compass at o Degree (or at the Equator,) and move the Point k to the Zenith Distance, (the Complement of the Altitude, taken by the usual Instruments,) and from any Hour, as from C, describe an Arch of a Circle with the faid Point, as bc (Ex. 1.) Suppose eight Hours after, by your Watch, you have another Observation: move the Meridian eight Hours farther, to d, and fix it there; and with the Zenith Distance then observed, describe another Arch as ef, the Point where it cuts the former is the Place of Observation, and its Distance taken on the Meridian from the Equator, shews its Latitude; and the Minutes reckoned on the Equator from the Meridian Meridian to C and d (the Times of Obser-

vation) shew what those Hours were.

Case II. When the Sun has Declination: Fix the Centre of the Beam-Compass on the Meridian, to the proper Degree of Declination for the Day of Observation, and

proceed as before.

Case III. If the Observations are at a greater Distance than twelve Hours, but in the same Day: Make use of the Complement to twenty-four Hours of the Distance in Time, and take the Declination on the contrary, or lower side of the Equator; and instead of the Zenith Distance, take the Nadir Distances or Altitudes increased by 90°.

Thus you will find the Latitude, and Time of each Observation from Midnight. In this Case the Beam-Compass must extend

to more than go".

Case IV. If the Observations are more than a Day asunder; as for Instance, a Day and 2 Hours (26 Hours): Place the Center of the Beam-Compass 2 Hours farther than it was the Day before; but in different Declinations, according to the Table of Declination for the several Days.

Case V. When the Observations are made by a Star: The Centre of the Beam-Compass must be set to the Declination of the Star; then proceed as before. To find the Hour in this Case, the right Ascension must

be likewise given.

Scholium.

G 2

[44]

Scholium. The same Method may be useful at Land, when no Meridian Observation offers.

II. The Ship in Motion.

Case I. Suppose the Sun in the Equator: The Distance between the two Observations 8 Hours, as before, and the Arch a a a (Ex. 2.) described by the Zenith Distance of the first Observation, from the Centre C; and the Angle cab, 40 Degrees, is the Angle between the Ship's way, and the Azimuth of the Sun continued, (given by the Azimuth Compass) and that during the eight Hours, the Ship has made 1°, or 60' from a to b, or from the Sun; then, as Radius is to the Cosine of cab 40°, so is ab 60' to ac 46'; add 46' to the Zenith Distance Ca; and with k, the Point of the Beam-Compass set at that Distance, describe the Arch cbe; then with the Zenith Distance of the last Observation, whose Centre is d, draw the Arch ff; the Point where it cuts the Arch cbe. is the Place where the Ship was last; and its Distance taken on the Meridian from the Equator shews its Latitude; the Minutes reckoned on the Equator from the Meridian to d (the Time of the last Observation) shew the Hour, or its Distance from 12 o'Clock.

Case II. If the Ship had sailed from a to β or towards the Sun: The Cosine of the Angle $\beta a \gamma$, or of the Angle between the Ship's

[45]

Ship's Way and the Sun, must be subtracted from the Zenith Distance of the first Observation.

N. B. Only the two Arches cbe, ff, are to be drawn on the Globe, the rest being added here, to shew the Reason of the Construction.

Case III. To find the Latitude of the first Place: From the Equator, with a Pair of Compasses, take the Distance sailed 60', and with one foot in the Intersection of the Arches be, ff, the Place sound before, put the other in the Arch a a a, the Zenith Distance of the first Observation, and in this Instance, on the lest Hand of the Azimuth of the Sun, this is the Place sought; and its Distance taken on the Meridian from the Equator, shews the Latitude; and the Minutes reckoned on the Equator from the Meridian to C, the Time of the first Observation, shew the Hour.

The Interval in Time or Degree between the two Places, shewn by the Index G, is the Difference of Longitude.

N. B. Those Observations are best, whose Arches cross each other almost at right Angles.

Prop. II. The Zenith Distances of two Stars, observed at the same Time, their Declination, and right Ascension being known;

[46] to find the Latitude of the Place of Obfervation.

Fix the Center of the Beam-Compass to the Declination of either of the Stars, and with the Zenith Distance of that Star describe an Arch; move the Meridian as many Hours farther as is the Difference of right Ascension of the other Star; and fix the Centre of the Beam-Compass to the Declination of it; and with its Zenith Distance cross the first Arch: The Intersection shews the Latitude of the Place of Observation; and also the Distance of the right Ascension of the Zenith from that of either of the Stars, by which means the Hour may be known.

If a Celestial Globe is made use of, then place the Centre of the Beam-Compass over

the feveral Stars.

The Latitude and Hour being given, the Variation of the Compass is easily known.

N. B. In order to draw Arches on the Globe; rub some black Lead powdered on a Piece of Paper; lay the Side which is blacked next the Globe, where you imagine the Intersection of the Arches will be: Then draw them on the clean side with the Point of the Beam-Compass, and they will appear on the Globe; and if the Globe is well varnished, they may be rubbed out with Bread, or washed out with Water.

F 47]

As Altitudes at Sea are now readily taken, with great Exactness, by Hadley's Quadrant, and as the said Altitudes are the Principles on which the Operations above described are founded; the previous Use of that Quadrant cannot but be of the utmost Importance to those who shall have Occasion for this Instrument; [for which, see my Theory of Hadley's Quadrant demonstrated.]

SECTION IV.

The USE of the GLOBE in finding the true Distance of the Moon from the Sun or a STAR, by the observed Distance, allowing for REFRACTION and PARALLAX.

** HE Globe is an Instrument equally applicable for finding the true Distance of the Moon from the Sun or Star, from the observed, or apparent Distance, taken by the SEA-OCTANT in the Problem of finding the LONGITUDE, or Difference of MERIDIANS, between any two Places. For the Degrees upon the Globe of 28 Inches Diameter, are larger than those upon

upon the Octants of the usual Size; and therefore by means of the Nonius, the true Distance may be had to as great an Accuracy, at least, as the observed Distance itself; and to pretend to greater Precision than the Data will admit of, is absurd; and will in vain be sought after; even from Calculations themselves.

To illustrate this Matter, we must explain by a Scheme the Doctrine of Refraction and Parallax, and the Difference between the Apparent and true Places of Objects, therefore (in Fig. 2.) let C be the Centre of the Earth AB; and B the Place of a Spectator on its Surface, whose Zenith is Z. And let N and n be two Places of a Planet, in an Azimuth Circle ZO; daw BN and CN, as also Bn and Cn; then are the Angles BNC, and BnC, the Parallaxes of the Planet at the Altitudes NO, nO, that is, if DG represent the Sphere of the Stars, a Spectator at the Center C, will view the Planet at N, among the Stars at F; and the Spectator on the Surface at B, will view it at E; and the Difference of these two Places E, F, in the Starry Firmament is the Parallax of the Planet, and is measured by the Angle BNC = ENC.

Now from the Principles of plain Trigonometry, we have BC: CN:: Sine of BNC: Sine of NBC = Sine of NBZ.

Also BC: Cn:: Sine of BnC: Sine of

CBn = Sine of nBZ, therefore (fince Cn = CN) we have the Sine of BNC: Sine of BnC: Sine of ZBN: Sine of ZBn.

But in finall Arches, or Angles at N and n, the Arches and the Sines are equal. Whence the Parallatic Angle at N is to that at n, as the Sine of the Planets Distance from the Zenith ZN, to the Sine of the Zenith Distance Zn.

Hence the following Corollies. (1.) If the Distance of the Object, or CN, be indefinitely great, the Parallax will be insenble, as in the Case of the Stars. (2.) If the Object be in the Zenith at Z, the Angle CNB, or the Parallax, vanishes. (3.) The Parallax is greatest of all when the Object is in the Horizon at M, 90 Degrees from the Zenith Z. (4.) The Horizontal Parallax BMC is greater or leffer as the Diftance CM becomes less or greater. (5.) The Horizontal Parallax at M is to the Parallax at any Altitude N, as Radius CM to the Sine NH of the Zenith Diftance. (6.) The Diftance of the Planet (CN) is to the Semi-diameter of the Earth (BC) as the Sine of the Zenith Distance, to the Sine of the Parallax, or as the Radius to the Horizontal Parallax. (7.) The apparent Zenith of the Planet among the Stars is GE and the True Zenith Distance is GF; always less than the apparent.

In like Manner may be explained the Ef-H fect fect of the Atmosphere in refracting the Rays of Light. Let AB (Fig. 3.) be the Surface of the Earth, C its Center, B the Place of a Spectator; M N Z the upper Part of the Atmosphere, F any Celestial Object; F N a Ray of Light falling upon the Atmosphere in N, but is by the encreasing Density of the Air, gradually refracted into the curved Line N B to the Eye of the Spectator; let B E touch the Curve in B, and E will be the apparent Place of the Object F among the Stars, by the Laws of Optics.

Hence it appears (1.) That the apparent Zenith Distance GE is always less than the True Distance GF. (2.) The Object in the Horizon at O, will be most elevated by Refraction. (3.) An Object at the Zenith G suffers no Refraction at all. See my New Elements of Optics for more upon this

Subject.

The Application of this Doctrine of Parallax and Refraction in finding the Longitude at Sea, by Means of the Celestial Globe of 28 Inches Diameter, will not be difficult. For this Purpose the Altitude or Zenith Distance, of the Moon, and of the Sun or Star at a given Instant, must be carefully taken with a Sea-Octant, as also the Distance between them; and from these Measures of the apparent Distances, and the given Quantities of Parallax and Refraction at the respective Altitudes, the true Distance and Places

of the Moon and of the Sun or Star may be found upon the Surface of the Globe, furnished with the following simple, and easy Apparatus, viz.

First; An artificial lunar ORBIT, which may be most commodiously constructed with a filken String or fine Cord, and properly

rectified to the present Month.

Secondly; Two moveable graduated vertical' Circles to be fixed to any two opposite Points of the general Meridian, or the Zenith and

Nadir of the Place of Observation.

Thirdly; A graduated Measuring Arch, furnished with a fixed and moveable Nonius' for measuring Distances to a Minute, with Ease. Then to represent the Process, let HZO (Fig. 4.) be the Brass Meridian, the Zenith; A the apparent Place of the Moon; EA her Altitude, or ZA the apparent Zenith Distance taken by the Qua-Also let B be a Star, whose apparent Altitude is BF, or Zenith Distance Z B; then of Course A B will be the apparent Distance between the Moon and Star, all measured as exactly as possible by the Quadrant at the fame Time.

Now by a Table of Refractions (which you find in my THEORY of HADLEY'S QUA-DRANT) the Refraction for the Altitude FB is known, which is B*, and to be fubducted from the apparent Altitude F B, and there will be had the true Altitude F * of the Star.

H 2

L 52

Star. The Place of the Star, truly corrected upon the Globe, will coincide with the Point * of true Altitude in some Position of the Vertical ZF, which is easily found upon Trial.

Then by the same Table the Refraction is found for the apparent Altitude of the Moon A E; but besides this, the Quantity of Parallax must also be sought for the said Altitude, (and which is always to be had in every Set of Astronomical Tables.) But as the Refraction encreases the Altitude, and the Parallax diminishes it, the Difference between both must be taken, and added to the apparent Altitude A F, and this will give the true

Altitude of the Moon EC.

Star

This being obtained, fet the Index of the fixed Nonius to the apparent Place of the Star B upon the Vertical ZF, and carry the meafuring Arch to the other Vertical, and move them both in such Manner, that the apparent Altitude E A, and measured Distance BA, may both coincide at A. This gives the true Position ZE of the Moon's Vertical; then, every Thing being rightly ordered, the true Place of the Moon) will fall upon its proper Orbit; and the true Distance between the Moon and Star, viz,)*, will be found for the Time of the Observation, by the measuring Arch, and its Nonius, to a Minute, If he apparent Alicude F P. stuniMis of the If the Distance of the Moon from the Sunbe taken, let that be AD; then pursuing the same Method as before, you get the true Distance of the Time and Place of Observation. As these apparent Distances are taken between the Limbs of the Luminaries, Allowance must be made for their Semi-diaters at that Time, that the true Distance of

their Centers may be known also.

Then for the same Moment of Time let the Place of the Sun and Moon be calculated for the Place you would know the Longitude from, as London, Paris, &c. By this Means you have the Difference of the Moon's Place in her Orbit, at those two (really different) Moments of Time; and then finding the Moon's Hourly Motion, you find thereby the Time required for the Moon to pass through that Arch of Difference, which Time is the Difference of Longitude sought, when converted into Degrees, as directed Page 42, of the Use of the Globes.

The Instrument for taking the Distance of the Moon from the Sun or Star, is Had-ley's Octant, the only one, indeed, adapted for such an Use, but as they are usually made, they can be applied only in the first and last Quarters of the Moon in measuring her Distance from the Sun; and her Distance from a Star more than 90 Degrees, cannot be taken. There is therefore a considerable De-

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fect, yet remaining in the Instrumental Part of this famous Problem.

It is to remedy this, that I have proposed to construct this Instrument with two Indexes, viz. one for Degrees, and the other for Minutes of a Degree; so that by Means of both, you measure any Arch or Angle to 180 Degrees, which is the greatest of all; and this without any Addition to the Bulk of the Instrument, which is rather less in this, than in the usual Form, and equally easy to be applied for measuring any Distance between the Moon and the Sun or Star.

I shall only observe further on this Head, that my New Goniometer may be applied to the Nautical Top, for taking Altitudes of the Sun at Sea, instead of Hadley's Quadrant; and thus the Multiple Angle of Altitude, will be true in Seconds, when the common Method (of measuring the Angle at

one) rarely gives it to a Minute.

For the Moon from the Sun or Star, is Modeller's Office, where or by over, and each adapted about not be they are usually under

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SECTION V.

or Marketter or under True 212,

The Rationale of the METHOD of correcting the Places of Stars, when necessary, upon the Celestial GLOBE: My Company of the Company of

*** T has been shewn (Page 100, &c.) 1 that by the Retrograde Motion of the Equinox, the Stars have all an apparent Motion in Consequentia, i. e. from West to East, at the Mean Rate of one Degree in 72 Years, or 50" per Ann. The Consequence of which is, that the Longitudes or true Places of the Stars will be constantly altering in the Heavens, and therefore cannot long agree with their Places upon the Surface of the Globe. Also their Right Ascension, Declination, Altitudes, Azimuths, Amplitudes, &c. will in a Course of Years be fenfibly variable; and therefore require to be rectified for the Uses of GEO-GRAPHY, ASTRONOMY, and NAVIGA-TION, &c. about every 20 Years. But

orl'T'

But before we give the Rules, we stall show how, and from whence they are derived; and the Reader may find the Demonstration of the Whole at large in Vol. II. of my Mathematical Institutes, Page 253, &c. Therefore let HZO, (Fig. 5) be the Meridian, Z the Zenith, HO the Horizon, P the Pole of the Ecliptic EQ, and N the Pole of the Equator ÆQ. Also let A be the Place of a Star, NAB a Circle of Declination, and PAC a Circle of Latitude.

Then in the Ecliptic, C will be the Star's Place or Longitude, and let its annual Variation be $Cc = 50^{\circ\prime\prime}$, through A draw e Aq parallel to the Ecliptic E Q, and draw the dotted Circle of Latitude, c P, cutting the Parallel e q in (a) which will be the Star's Place in its Parallel after one Year, and its Change of Place, or Motion in Consequentia, will be A a dougle of the parallel after one Year, and its Change of Place, or Motion in Consequentia, will be

Again, in the Equator Æ Q, the Alteration in the Star's Right Ascension will be Bb; and that of its Declination will be A B—ab. Now as A and a may be considered as two different Stars in the same Parallel of Latitude cq, 'tis evitlent what ever relates to the Altitude, Azimuth, Amplitude, Hour-Angle, &c. of two such Stars, will be the same in regard to the two different Places A and a of the same Star, by its Motion in Consequentia.

The

The Longitude and Latitude of most of the noted Stars, especially those about the Zodiac, the late Abbe De la Caille has given us, in his Fundamenta Astronomiæ, corrected to the Year 1750. By which it will be easy to find the Longitude, Declination, and Right Ascension, of any of those Stars, and consequently their true Places upon the Globe, for any Year proposed, by the following Rules and Examples.

Let the Longitude of Sirius be required for the Beginning of the Year 1767. Then in the Table the Longitude of that Star was in \$5, 10° 38′ 22″, at the Beginning of the Year 1750; and in 17 Years, the Motion in Longitude is 17 × 50″ = 850″ = 14′ 10″ which added to the above Longitude, makes \$5 10° 52′ 32″ for its Longitude re-

quired.

man's

To find the annual Alteration in Declination of any given Star, as Regulus, or the Lion's Heart; the Analogy to be used, requires the Star's Right Ascension from the Solstitial Colure, or Angle ÆNB to be known. Now the Right Ascension of Regulus being 148° 51' 6", if you deduct 90°, there will remain 58° 51' 6" = Angle at N; and putting the Obliquity of the Ecliptic = 23° 29', say

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As

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To the Alteration in Declination in policie Vi 21318 33

From hence it appears, that the Variation in Declinations is proportioned to the sine of the Star's Distance from the Solstitial Colure only intherefore those Stars in oninedry that Colure, have the fame Declination for many Years together; but those in bnaneabethe Equinoctial Colure, fuffer the greatest Change ing to the following Specimen noisenilose To To find the Annual Alteration in Right Ascension, the Angle at A must be known; and for that Purpose in the oblique Triangle ANP, there is given (for Regulus) AN = 76° 49', the Co-Declination; the Angle at P = 56° 21', the Longitude from the Co-lure; and PN = 23° 29' the Obliquity of the Ecliptic; to find the Angle at A. fay SIRIUS As the Sine of AN = 76° 49" - 91988490 To the Sine of the Angle at $P = 50^{\circ}$ 21 $\frac{10.031}{9.92}$ So is the Sine of $\frac{1}{2}$ $\frac{1}{2}$ $\frac{10.031}{9.60}$ 10,000 18 88 08 7 12 7 30 38 34 26N The Eact | 294 46 72 44 To the Sine of the Angle A = 19 55 1-4 9.534360

Then

Then to find	B 6;	the Ale	eration	in Right
Ascension, say,				To the Sine

As the Sine of ____ AN = 76° 49′ = 9.988401

To the Co-tangent of $A = 19^{\circ} 55' - 10.440903$ So is the Variation in Declination 17" - 1,231835

8872731 Hence it appears that the Variation

76 the Vallation in Right Alcention 484 1977 in 1684 337

the true Places of the Star for any Year propoled mand thus they have been found and tabulated for all the principal Stars, according to the following Specimen for STIARS of the first Magnitude, adjusted to the Year 1760; Nessmin A to also A Principal H.

Names.	Right	Alcen	lion.	Ann.	as D	eclin 9/80	tion.	Affn.
he Angle as	65	istin	hils	114	00	514	200	1º9
CAPITAL A STI	74	44	53	66	45	43	34N	†
A ORION	75	報	23	43	14	29 20	46S	04 V
SIRIUS PROCYON	98	38	45	40	16	24	5S	ayt
RECORUS "	148	53	28	48	5	49	29IN enoN	ti ak
ARCTURUS!	198	-8	47	47	9	54	IS	+14
ANTARES	243	41	4	55	25	52	368	40
The LYRE The EAGLE	277	12	. 7	30	38	34	26N 9N	+
FOMAHAUT?	341	5	3	50	30	53	125	-1

It is evident from the Figure, that in the first and third Quarters of the Ecliptic, the North Declinations energie, and their Variations are affirmative, or to be added; but in the second and fourth Quarters, they are Negative, and to be subtracted; and vice versa in Stars of a Southern Declination.

The retrograde Motion of the Equinex can make no Alteration in the Latitude of Stars, because the Equinoctial Point is always in the Ecliptic, and consequently the Motion in Consequentia, in every Star, is in a Parallel to the Ecliptic. But it happens, that on Account of the attracting Forces of the Planets, the Ecliptic itself is not stable. of constantly in the same Position with respect to the Equator, and therefore neither the Obliquity of the Ecliptic, not the Betitude of Stors can continue presifely the same; yet the Variation in 100 Years does not amount to any septible Quantity (viz: about 40 Seconds more or less) and therefore can make no Alteration of their Blaces upon the fing in the Morning; and it let Het Bolon

when it ceases to be feen in the Honzon of an Evening after Sum-let. It has been found by Ptolemy, and other Astronomores, that the Distance of the bun below the Horizon, is then about 12 Degrees, for a Star of the first Magnitude.

T Pherefore to find the Hehacal ribing of a Star, bring it to the eaftern Part of the Hori-

It is evident from the Figure, that in the said and third Quarters of the Ecliptic, the secure was a said was a said on the Ecliptic, the said on the second and fourth Quarters, they are North to be said on the second and fourth Quarters, they are not seen and sold of the second as seco

The pacient Canicular Astronomy
to explained upon its proper Prinle ciples, and illustrated by Calcuand vine position of the page of t

TAX THE HELTACAL RISING and SET-A STING of the Stin was omitted paramong the Problems of the Stars. distributione finally liere fupply it; especially as the Doctrine of the Cancular ASTRONOMY (concerning which the learned BATNERT DGE wrote a Treatile) depends entirely upon it. TUA Star is faid to He Heliacally, when it is ate a fufficient Distance from the Sun to be fielt feen diffindly in the Hotizon at its Rifing in the Morning; and it fets Heliacally. when it ceases to be seen in the Horizon of an Evening after Sun-fet. It has been found by Ptolemy, and other Astronomers, that the Distance of the Sun below the Horizon, is then about 12 Degrees, for a Star of the first Magnitude.

Therefore to find the Heliacal rifing of a Star, bring it to the eastern Part of the Hori-

the wellern Part of the Globe, walld therever move it to that the fight Degree may cot they Ecliptic; mark that Point in the Ecliptic, and the Point opposite to it, is that in which the Sun is 12 below the Horizon, when the Star rifes, and the Time of the Year corresponding thereto, is found in the Calendar upon the Horizon. In the same Manner your find the Time when the Star sets believely.

It is supposed, the ancient Egyptians took this Method of observing the Moon, in order to get some Idea or general Measure for Months; hence, they concluded, that the Time between two beliatal Risings of the Moon (or between two New Moons, as they are now called) was 30 Days, and accordingly they divided their Year into 12 equal Months of 30 Days each; this was called the Luni-Solar Year; and confisted of 300 Days.

But having learned by continued Observaltions on the Heliacal Fifting and setting of Stars that the true Length of the Year exceeded 360 Days, by about 5 Days, they added these 5 Days to the End of the Year, which then consisted of 365. These additional Days were afterwards called Boase a menæ by the Greeksanin A sti to (2012) 1613

Because the Egyptians in these Oblavair tions made use of the Britis dilled the Star) this Year of 365 Days was called the [630]7

Canicular Year. As Astronomy farther advanced, it was soon abserved that this Year was deficient of the Length of the natural Year by about 6 Hours, or a Quarter of a Day defor by that Space of Time they found the heliocal Rising of Sirius was later and languaged yearly. Therefore in 4 Years the Star rose one Day later, and in 4 times 365, or 1460 Years it rose a whole Year later, that is, in the Compass of 1460 Years, Sirius would rise beliacally upon the same Day of the Month.

This Space of Time was by the Egyptians called the Sothiacal Period, because the Name of Sirius, with them was $\Sigma \omega \theta$, Soth. And of Course, 1460 Solar Years were equal to 1460 of the Egyptian Civil Years; the first Month of which was called Thath, and Corresponded to Saptember in the Roman Calendar, And it is to be supposed that the Beginning of the Sothiacal Period was at the Heliogal Rising of Sirius on the first Day of Thoth, or the Beginning of the Egyptian Year as Yea

YMr. BAINBRIDGE illustrates this Matter by the following Scheme, (Fig. 6.) Let SZ he an Asch of the eastern Horizon, OZ a Part of the Ecliptic, and S the Place of the Star (Sirius) at its Rifing. Let the Sun be rifing at the same Time in Z, by his Light objecting the Star at S. Now that the Star may be will be at Rifing. it is required that

the Sun be at some certain Distance from the Star, below the Horizon, and let that Distance be Z a: so that when the Sun is in a or lower. the Star rifing at S will be visible, but if the Sun be any-where between Z and a, it will

not appear at rifing.

From a towards Z, take the Arch as equal to a Quarter of a Degree; and on the other Side, take $\alpha \delta = \beta \gamma = \gamma \delta = \alpha \epsilon$; and the Motion of the Sun through a Quarter of a Degree is made in 6 Hours nearly. which is the Excess of the Tropical Year above the Egyptian Year of 365 Days.

These Things premised, it follows, that on the first Day of Thoth, in the first Year, SIRIUS at S is rifing and visible when the Sun is at &; the fecond Year, on the same Day of Thoth, the Sun will be at y; in the third Year, on the first of Thoth, when he is at B; and the fourth Year still on the first Day of Thoth, when he is at a; for at the End of every Egyptian Year, there was wanting to compleat the tropical Year, a Quarter of a Day; and to finish the Sun's Course, a Quarter of a Degree.

During these first 4 Years, then, when the Son is in 8, 2, B, a, the Star rising at I will be visible on the first Day of Thoth. But in the 5th Year, on the first Day of Thoth, the Sun will be at e, and the Star rifing at S, will not be feen; but on the fecond Day of Thoth, the Sun will be in &,

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the first Fear, on the same second Day of Thotby the Sun will be in γ ; the seventh Year in β ; and the eighth Year in α . So that during these second four Years, the Star S will rise Heliacally visible on the second Day of Thoth. And after the same Manner it is shewn, that during the next four Years, the Star will rise and be visible on the third Day; and for the next four Years, on the 4th Day of Thoth; and so on, till it has passed through all the Days of the Year.

Having thus explained the Nature of the Canicular Year, and Sothiacal Period; Mr. BAINBRIDGE next proceeds to an aftronomical Solution of the following Problem, viz. For any given Time and Latitude of a Place, to find the Heliacal Rifing of any given Star.

In order to the Solution of this Problem, the Longitude and Latitude of the Star, as also its Right Ascension and Declination, are to be taken from Astronomical Tables; from thence we find the Ascension; and these being known, we find the Point of the Ecliptic rising with the Star, and the Angle contained between the Ecliptic and Horizon; and then, lastly, by these Pracognita, we find the Sun's Place in the Ecliptic, in the beliacal Rising of the Star.

Having confirmated a proper Scheme, (Fig. 7.) for Heliopolis in lower Egypt, we can thereby

chereby illustrate the Process of astronomical Calculation; for therein let HZOQ be the Meridian; HCO the Horizon; P the North Pole; OP the Elevation thereof, or Latitude of Heliopolis; Z the Zenith. ÆCQ the Equator; EAL the Ecliptic; at the Autumnal Equinox; S the Place of Sirius in the Horizon rising heliacally; PSM a Circle of Declination; ISR a Circle of Latitude; N the South Pole of the Ecliptic. ZST the vertical Circle through Sirius. D the Place of the Sun at the heliacal Rising of the Star, to find which is the general Problem proposed.

Therefore let the Example be to find the beliacal Rifing of SIRIUS at Heliopolis in the Year 1750. The Latitude of Heliopolis is $EZ = 30^{\circ} 22'$; and the Declination of Sirius, is $SF = 10^{\circ} 23' \frac{1}{4}$; therefore in the Right angled Triangle SFC, say (by Case II. Page 192 of the Use of the GLOBES.)

As the Tang. of the Angle SCF = 59° 38' - 10.232165

Is to the Tangent of SF = 16 23 - 9.468580
So is Radius 90 10.

To the Sine of the Arch CF = 9 47 - 9.236415

To the Right Ascension of Sirius for 1750, viz. 98° 32', add the Ascenson Disference now found, and the Sum will be 108°: 19' = oblique Ascension of Sirius.

The

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The next Thing to be determined is the Point of the Ecliptic A rising with Sirius S, to find which there is given in the oblique Triangle C - A, the Obliquity of the Ecliptic or Angle at $c = 23^{\circ} 28'$; the Angle $c = 23^{\circ} 38'$, the Co-latitude; and $c = 23^{\circ} 38'$, the Co-latitude; and $c = 23^{\circ} 38'$, the Supplement of Oblique Ascension (just now found) $c = 23^{\circ} 38'$, to find the Side $c = 23^{\circ} 38'$.

In order to this, from the Point C let fall the Perpendicular Ga, then in the right angled Triangle C = a there are known the Hypothenuse C = a and the Angle C = a to find C = a; say (by Case III.)

As Radius	10.
To the Sine of $\triangle C =$	710 41'- 9.977419
So is the Sine of C $\triangle a =$	23 28 - 9.600118
To the Side C a=	Section 1
To the side C 2 =	22 12 - 9.577537

Then you find the Angle $\triangle C a$ (by Case III. Anal. 3.)

As Co-fine of $\triangle C =$	710 41' - 9.497301
Is to Radius So is the Co-tangent of	10. 23 28' — 10.362389
To the Tangent of $\triangle Ca =$	82 14' - 10.865088

Then from 82° 14' take 59° 38', there will remain the Angle A C $a = 22^{\circ}$ 36'.

In the right angled Triangle $\triangle Ca$ we find the Base $\triangle a$, by saying,

what!

2 A

As Radius — — 10.

To the Sine of $\triangle C =$ — 71° 41′ — 9.977419

So is the Sine of $\triangle C =$ — 82 14 — 9.995998

To the Sine of $\triangle a =$ — 70 9 — 9.973417

In the right angled Triangle A Ca, there are known the Side Ca, and the Angle A Ca; to find the Side Aa (by Cafe I. Anal 1.) fay,

As Radius
To the Sine of C a = -So is the Tangent of A C a = -22° 12' - 9.577537
22 36 - 9.619364

To the Tangent of A a = -8 56 - 9.196901

Then from 70° 9' (= 2 a) take 8° 56', there will remain 61° 13' = 2 A; which taken from 180° , will leave 128° 47', or 8° 47', for the Point of the Ecliptic A rifing with Sirius at S.

To find the Angle at A, fay,
As the Sine of A $a = -\frac{8^{\circ}}{5^{\circ}} = \frac{9.191130}{9.584665}$ To the Sine of A $Ca = -\frac{22}{2} = \frac{36}{9.577537}$ To the Sine of C $Aa = -\frac{22}{12} = \frac{12}{9.577537}$ To the Sine of C $Aa = -\frac{69}{19} = \frac{9.971072}{9.971072}$

Lastly; in the right angled Triangle ABD, there are known the Angle BAD, last found, and the Side BD, the Depression of the Sun below the Horizon, which to render

render the very large and bright Star Sirius vifible at S, in the clear and serene Air of Egypt, Mr. Bainbridge estimates at 11 Degrees.

Therefore fay,

As the Sine of BAD = _____ 69° 19′ ____ 9.971072

To the Sine of BD = _____ 11 00 ____ 9.280599
So is Radius _____ 10.

To the Sine of AD = _____ 11° 46′ ____ 9.309527

To the Co-orient Point of the Ecliptic A

= \$1.8° 47' add 11° 46', the Sum will be
\$1.20° 33' for the Place of the Sun required,
when Sirius rose heliacally in the Year 1750,
which was on Aug. 3d. Old Style.

Now from Aug. 3. to Sept. 1. (or Thoth) when the Sothiacal Period begins, is 28 Days; therefore 28 × 4 = 112 Years yet to come of the Current Period; hence 1460 — 112 = 1348, the Year of the Period for An. Dom. 1750. Also 1750 — 1348 = 402 the Year of Christ, when the present Period Perio

riod began. Again 1460 — 402 = 1058 the Year before CHRIST when the former Period began. And lastly 1750 + 112 = 1862, when the Sothiac Period begins again.

Hence it appears that the learned Bainbridge was mistaken in fixing the first Year of the Sothiac Period to the Year of Christ 138, by following the uncertain Accounts of Censorinus and Clemens Alexandrinus. Also Mr. Costard has committed an Error in taking [70]

taking the Arch AG of 57° 34' inflead of 27° 34', and thence makes the prefent Sothiacal Period begin A. D. 375 which is 27 Years too early. See his first LETTER to MARTIN FOLKES, Efq; Page 34, &c. And Bainbridgii Canicularia, Page 34, 35 The Substance of his Book we have now nearly exhausted, and shall only further obferve that when Sirius rifes beliacally, those called the Dog-Days end, and they begin when this Star fets beliacally, i. e. when he is less than 12 Degrees from the Sun at his Setting; for during that Interval, this Star is above the Earth with the Sun, and because when the beliacal Rifing and Setting were in the Beginning of the Month Thoth, thefe Dog-Days were very hot, it gave Rife to the superstitious Notions the Ancients had of the pernicious Influence of this Star on the Air. the Bodies of Animals, &c. But as they had no Intercalary Day, their Dog-Days in one Sothiac Period run through all the Months of the Year, whereas, with us, by Means of the Day added in Leap Year they are always in the Autumnal Season. when the Salary Period In the ser

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SECTION VIL

An Abstract of Mr. Hornsby's Account of the ensuing Transit of Venus, in 1769, and the best Places for observing the same; to which is subjoined an Account of a New Helioscope to be used for shewing the Phases of the Transit to the greatest Advantage.

*** N my TREATISE of the TRANSIT I of VENUS over the SUN's Difk, printed in the Year 1761, the Reader will find most of the Particulars relative to the THEORY and CALCU-LATION of fuch a Phanomenon; and whereas the Observations on the Transit of Venus in that Year, were not of much Use in determining the SUN's PARALLAX; it is to be prefumed that the Transit of that Planet over the solar Disk in the Year 1769, will be thought worth a more mature and thorough Attention; and though Literature feems now to be in a manifest Decline, yet as such a Means of determining the Solar Parallax, and in Consequence of that, the Distances of all

the Planets from the Sun, and from each other, will not offer again in more than an 100 Years, there is no Doubt but such a fortunate Opportunity will be eagerly embraced by all who have it in their Power to contribute any ways towards improving the same; and indeed it is no less than a Duty the prefent Age owes to Posterity, who must think themselves greatly disregarded, if an Age abounding with such Riches, learned Men, most accurate Instruments, and in short, every Thing necessary to ascertain the Dimensions of the Solar System, from two Transits (offered as it were on Purpose) were not to do this great Work for them. need not mention what an eternal Difgrace it must reflect upon the present Generation, should they supinely let such a double Opportunity flip for promoting the Honour of the Sciences in general; and in particular for advancing ASTRONOMY to the Apex of all its Glory.

As the Problem of exhibiting all the Phases of such a Transit, is one of the most noble and useful upon the Globe, I have been very circumstantial in giving Directions about it in the Treatise of the Globes, Page 140, &c. exemplifying the same for the said Transit in June, 1761; and also by a large Print of the same, to give the best Assistance I could for forming an Idea of such an interesting Assistance. As these Problems will be the same for

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any Transit; it will be only necessary here, to state the Times of the Phases, and Duration of the Transit in 1709, which we have already done to our Hands by Mr. Hornsby, in a general Manner, from Dr. Halley's Tables

of the Motions of Venus.

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In Fig. 8 ABD is the Sun's Disk, C its Center, AN the Ecliptic, EN the apparent Path of Venus, K is the Place of Venus at the external Contact with the Sun's Eastern Limb, H her Place at the internal Contact with the same, H I her Passage over the Solar Disk, F her Place at the true Conjunction, and G at the Middle Moment of the Transit; then the Particulars of the Transit are specified by the solar lowing Numbers.

10Wing Numbers.	Survey to and
The Mean Longitude of Venus in the Beginning of the Year 1769	0 5 23 48
The Place of the ascending Node	2 14 35 21
The Ecliptic Conjunction, June Venus's Geocentric Latitude CF	d. h. / // 3 9 59 24 10 135
Logarithm of the Earth's Distance from	35.0065166
Ditto of Venus from the Sun	4.8610947
Ditto of Venus from the Earth	4.4606784
Equation of the Precession of the Equino Venus's hourly Motion from the Sun the Ecliptic	in { 3' 57.7"
Ditto in the relative Orbit F. N -	1'1 000"

es trache Barth's Center. I o the British
The Angle of the Path with the Enliptic; 25829 2 wiz. AN Enliptic To form of the Enliptic With the Equator 7. 254
Angle of Venus's Path with the Equator T 15 32156
Leaf Diffance of the Centre Contact In 10, 6,8%
Mean Time of the Middle of the Transit w nuc adt at Greenwich Observatory at Greenwich Observatory at Creenwich Observation in Longitude 1 na man and 2551
Mean Time of the Middle of the Transit of Shut IA at G equated of the Transit of The 139 58
Aberration in Latitude 1,35 Equated mean Distance of the Centers C G 10 8,15"
The Semi-diameter of the Sin 2 of 2 o
Interval of Time between the external Control of Market at K, and Middle at God Sand Sand Sand Sand Sand Sand Sand San
Duration of the Ingress of the Start of the
Time of external Contact, June — 3 7 12 56 Of total Ingress — 7 31 52 Middle of the Transit — 10 23 4 Beginning of Egress — 13 14 16 Last Contact
Lan Contact The dead Inoth the about 113 33 11

Such would be the Phænomena of the Transit, without regard to Parallax, or viewed ed from the Earth's Center. To the British Isles, and to the neighbouring Parts of the Continent, the Effect of Parallax is nearly at a Maximum; and will considerably accelerate the Times of the external Contact and Ingress. If the Sun's Parallax be supposed 8,7" the horizontal Parallax of Venus from the Sun will be 21,87". Whence the Times of the external Contact will be at 74 5' 47" more than an Hour before Sun-set, when the Altitude of Venus is about 8 Degrees. And the total Ingress will happen at 7' 24' 40", the Height of the Planet 5 Degrees.

If the Sun's Parallax should be one Second larger, or 9.7", the Time of Ingress will happen at 7^h 23' 51". And for the Times above-mentioned the Phases of the Transit may be universally exhibited by the Globe in the Manner directed in the Body of the Book, Page 140, &c. for the Sun's Declination

22° 26' 40", at that Time, and seewing on the

It is farther observed by Mr. Hornsby, that the joint Effect of the Parallaxes in Longitude and Latitude to lengthen the total Duration of the Transit will be greatest to those Places which are about 24° or 25° to the East of Greenwich, and in the ooth or 67th Degree of North Latitude, when the Sun's Altitude at each Contact is about 5 Degrees; as at Tornea°, Kittis, and the adjoining Parts of Swedish Lapland. But if the Sun's Altitude be required = 101 the

Latitude of Places under the fame Meridian will be 930 broggerias at Mardban, Sec.

Ata Torner when first minternal Contact gives 6' 3" and the fecond 4047 and the Sum 11 40 is the Time by which the total Duration will be lengthened thereis flattent

On the other Hand, there are Places where the faid Duration will be morrened as much as possible , and the Latitude of these Places are 34 South; and im 123 of West Longitude from Greenwich, when the Sun's Altitude is 18 80 devint about 49 of South Latitude, when the Sun's Altitude is 10°.

On the Supposition that the Sun's Parallax is # 8 7" the total Duration of the Transit in Latitude 55 will be shortened no less than 102 53 ... The whole Difference therefore of the Duration of the Transituin this Latitude, and at Torteap is no less than 24 35" A Difference confiderably greater than was expected by Dr. HALLE win the Transit of 1761 . Soul Dide Land in state

Another very good Method of finding the Son's Parallax at this Transit, is, by flationing two Observers in such a Manner, what one of the internal Contacts might be observed with the greatest Difference possible arising from a contrary Effect of Parallax at the two Places. This Method may be practifed at both Contacts in 1769, but in this Way, the Longitude of the two Plates most be rigotoully known superied souling it and on

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then Greenwich and Dublin will be very advantageous Stations for observing the first internal Contact; and in South Latitude 46°, and West Longitude about 108°, will be the best Station for the last internal Contact, when the Sun's Height is 5°. But as these there does appear to be any Land, the nearest Hands must be made Choice of for that Purpose and when the sun's height is 5°.

The last internal Contact or Beginning of Egress will be accelerated as much as possible, when the Sun has 3 Altitude, in about 19° of South Latitude, and 123° of West Longitude from London. And under the Tropic of Cancer, and in about 67° of East Longitude from London, the said Beginning of Egress will be most of all retarded. But as many Degrees Difference in Longitude and Latitude will not make much Difference in Parallatic Time, so a large Scope will be left for chusing Places most advantageous for this Purpose.

As both these Methods may be practised in many Places, they will mutually confirm and illustrate each other. Thus far from

Mr. Hornefby: 01 the land the many of the con-

As to the Instruments for observing the Phases of the Transit, they are of the Telescopic Sort; but those which shew them by reslected Light without any Refraction thro'

Glaffes

Glasses at all, are most certainly the best; this may be done by a resecting Telescope by two Resections, and the Image of the Sun and the Planet will be extremely distinct, and the external and internal Contacts may be observed very critically without the Fatigue of poring through the Telescope as in the common Way of looking at the Sun, where a dark Glass must be used, besides the two Eye Glasses, all which tend to render these Observations less perfect and accurate.

The above Method when the Sun has about 5° of Altitude will do very well, as it makes the Sun's Image from fix to twelve Inches diameter, or more (if required) when the Room is large. But a HELIOSCOPE may be constructed for this Purpose, which shall give an Image of the Sun, by reflected Light only, of 24 Inches Diameter, and that of the Planet Venus 2 of an Inch, in so fmall a Space as 3 Feet by 21 ... This curious Contrivance I often use for viewing the Sun's Face and Spots, as also for shewing fmall Objects in vaftly large Images, by reflected Light, and therefore without Colours. Such a Method could want no Recommendation to the Curious, of they were but acquainted with its Merit; for which they may confult my OPTICAL Essays Page 30. The HEDIOSCOP E herewhentioned being only one of the many Lifes of the Univerfal Perspective there described. ole of the id nat Parts for Meate Mcba I do A



of poring through the Telescope as in the common Way of looking at the Sun, where

The Glaffes all which tend to render these



Should not have thought it worth my While to take Notice of any Thing faid by the Critical Reviewers relative to my Writings, but that I think a few Lines will be very necessary to apprize the Public how far they may think they have any Reason to be biassed in Favour of, or prejudiced against any Book, by any of their Criticisms upon it.

A few Years ago, a worthy and learned Trient of mine died, and his Widow published a Treatife of some very currents Mathematical Subjects contained in XXV Propositions, only was faulty, and wholly so; but the Critical Resources purposing to serve the Widow and Orphani, as far as lay in their Power, first employed a great deal of parrotic Eloquence to recommend the Performance, and to enforce it, descated very sagely on the Novelsy, Ingenuity, and Utility of the Tangential Sea Chart proposed by our Author in the XXIst Proposition above referred to. Now the Construction of this Chart was so widely different from that of Mercater, and every Way so false and unfit for any Naurical Turpole whatever, that it may fairly be presumed not a Tyro in Naurica, or a single Person in the Nation besides the Reviewer, who makes the least Pretence to any Skill in the Theory of Navigation, could have made so scandalous a Blunder as publicly to approve, applied, and recommend the Book on the supposed Truth and Excellency of that Invention!

Such is the Tribunal, at which the Merit of Authors is to be tried and decided. These are the Judges in Nautical Science, who reassured my New Paincipa, es, of Grography and Navigation, founded upon the New discovered Figure of the Earth, only because I could not tell which came search the Truth, the Charts constructed upon the Numbers resulting from the Degree measured under the Artic Circle as given us by Maupersuic, or the Tables of Dos Juans, derived from the Measure of a Degree under the Equator; and therefore my great Crime was to give them both, that the practich Navigator might have in his Power to the raine as he pleased, as they differ exceeding little from each other, and yet each comes very near the Truth, at least much nearer than the common Table of Meridianal Parts, or Mercator's Charts do.

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POSTSCRIPT.

These are the Connoisscurs who now undertake to inform the Publick what Opinion they should entertain of my Institutions of astronomical Calculations, but I believe the Mathematical Part of the Publick are too well acquainted with, and too good Judges both of their Writings and mine, to think themselves in the least obliged to them for their profered Service. This Work, they inform the World, is a Translation, which I did not know before; I never called it any Thing more than an English Edition of Mr. Claikaut's Tables of the Lunar Motions, with an Explication prefixed.

The 13th and 14th Chapters, they advice me to expunge in a second Edition, and why? because I have made Use of absurd Properties of Circles and Ellipses instead of the Radius of Curvature. I suppose this to be the first Time the Public ever heard that Circles and Ellipses had any absurd Properties———If they have none, I have made Use of none.

The Truth is no more nor less than this, my Design was to shew the true Quantity of the Phases of the solar Eclipse and Dimensions of the hunar Shadow upon the Earth's Surface, considered as an Ellipsoid, and not a Sphere; in order to which, I pursued the most direct, concise, and genuine Method that Geometry could supply for that Purpose; and this very Method of Demonstration was, some years ago, invented by the GREATEST MATHEMATICIAN in EUROPE, the universally celebrated Mr. de Maupertuis in his Treatise sur la PARALLAX de la Lune; and the same Principles and Method of Reasoning which he there employs for obtaining his Dimensions Geographique (Sect. III.); pour la Gravité (Sect. IV.); Recherche de la Difference des Parallaxes sur la Terre & sur le Glabe (Sect. VII.); Loxodromiques (Sect. XII); are the very same with those I have here used in these two Chapters; and no one that was not very dark indeed, could help seeing how ridiculous it must be to introduce the Radius of Curvature for such Purposes.

But what gives a Diffelish for these two Chapters, is, that they are, it seems, of my own Growth; this they say, but for what Reason I know not. I never regarded them as of my own Growth, being too conscious I was never able to grow any Thing so good. I thought every Body had known that this ingenious Method of Demonstration was the Invention of MAUPERTUIS, as much as that of prime and ultimate Ratios was of the great NEWTON.

I am therefore in no Humour to accept the Advice of these Nameless Gentry for rejecting that Method of Demonstration which I have received from the greatest Names in the Mathematical World. As to what they say of Mr. Simpson's Series being more simple, and better to calculate from than Mr. CLAIRAUT'S, I believe it will appear to be an errant Falshood to any one who shall compare them; for they will find that both the Series are the same, and Mr. Simpson positively declares they are so, in the Preface to his Miscellaneous Tracts.

From these Anecdotes, the Public will be very easily apprised how far they can conside in the Criticisms and Assertions of such Reviewers.

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